

Secondary Au–Au and Sn–Sn Interactions in the Superstructure of YbAuSn—Missing Link in the Series of KHg_2 Superstructures

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The stannide YbAuSn has been synthesized in quantitative yield by reacting the elements in sealed tantalum tubes in a high-frequency furnace. The structure was determined from single-crystal X-ray data: *Imm*2, $a = 469.7(1)$ pm, $b = 2191.2(4)$ pm, $c = 812.7(1)$ pm, $wR2 = 0.111$, $2061 F^2$ values, and 58 variables. It crystallizes with a pronounced KHg_2 type subcell ($a, 1/3b, c, \text{Imma}$). In the superstructure the tripled b axis allows an ordered stacking of Au_3Sn_3 hexagons with weak gold–tin, gold–gold, and tin–tin interlayer bonding interactions. The determined stacking sequence $- + - - + -, - + - - + -$ in YbAuSn realizes a so far missing sequence within a general ordering scheme of KHg_2 type superstructures. The group–subgroup relation in going from the KHg_2 subcell and chemical bonding in YbAuSn are discussed. Each ytterbium site has an ordered near-neighbor environment of six gold and six tin atoms in the form of two tilted hexagons. Magnetic susceptibility measurements show a nonmagnetic ground state of the ytterbium atoms. YbAuSn is a metallic conductor with a specific resistivity of $50 \mu\Omega \text{ cm}$ at room temperature. ^{119}Sn Mössbauer spectroscopic data show one signal at an isomer shift of $1.922(5)$ mm/s subjected to quadrupole splitting of $0.929(9)$ mm/s.

Introduction

The equiatomic gold stannides RAuSn ($R =$ alkaline earth or rare earth metal) exhibit an unusually large structural variety. With the trivalent rare earth metals we observe the hexagonal NdPtSb type ($R = \text{Y, Ce–Nd, Sm, Gd–Ho}$) and the cubic MgAgAs type ($R = \text{Sc, Ho, Er, Tm, Lu}$) structures.^{1–3} Among these stannides, HoAuSn is dimorphic.¹ The volume per formula unit is smaller for the NdPtSb modification (67.9 \AA^3) than for the MgAgAs modification (72.7 \AA^3). Detailed investigations of the magnetic properties^{4–14} clearly revealed the trivalent oxidation state for the rare earth atoms in

these intermetallics. The $[\text{AuSn}]$ polyanions of these compounds consist of slightly puckered, two-dimensional Au_3Sn_3 hexagons (NdPtSb type) or a three-dimensional network of corner-sharing $\text{AuSn}_{4/4}$ tetrahedra (MgAgAs type). No Au–Au nor Sn–Sn interactions occur in these polyanions.

This situation becomes totally different when an alkaline earth or a potentially divalent rare earth metal is introduced. So far, the structures of MgAuSn ,^{2,15} CaAuSn ,¹⁶ and EuAuSn ¹⁷ have been solved. While the cubic MgAgAs type is still formed with the small magnesium atoms, CaAuSn ¹⁶ and EuAuSn ¹⁷ adopt two different complex superstructures of the KHg_2 type. In both cases weak superstructure reflections forced a quintupling of the unit cells. The superstructures result from a different coloring of the gold and tin atoms on

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the subcell mercury positions. In both structures we observe puckered Au_3Sn_3 hexagons that are stacked in a different manner, resulting in interlayer Au–Sn as well as Au–Au and Sn–Sn bonding. The formation of the Au–Au interactions (relativistic nature of the gold atoms)¹⁸ and the tin dumb-bells is certainly a driving force for the formation of the superstructures. Also SrAuSn^{19} and BaAuSn^{20} crystallize with pronounced KHg_2 subcells. The work on these superstructures is in progress.

We recently developed a general group–subgroup scheme for the classification of different KHg_2 superstructures (ATX; A = alkaline earth, rare earth; T = transition metal; X = main group element).^{16,21} From a crystal chemical point of view the various KHg_2 superstructures can be categorized on the basis of stacking sequences of transition metal filled trigonal prisms formed by the X atoms. The scheme allows one to readily precalculate and predict superstructures of the KHg_2 type, i.e., the space group and the atomic coordinates. Helpful is that some of the KHg_2 type superstructures can be divided into two groups: (i) body-centered groups based upon $\text{Imm}2$ (EuAuGe^{22} derivative) and (ii) primitive groups based upon Pnma (TiNiSi^{23} derivative). With other words all EuAuGe variants are noncentrosymmetric and body-centered, while the TiNiSi variants are primitive and centrosymmetric. This limitation into two groups has its cause in the requirement of a reasonable crystal chemistry. The group–subgroup relations itself naturally allow an infinite number of superstructures, but taking the available data this is so far limited to only isomorphous (i) symmetry reductions.

The crystal chemistry expresses itself by stacking transition metal filled trigonal prisms and introducing weak interlayer bonding interactions between gold–gold, tin–tin, and gold–tin contacts, thus fine-tuning the structure. The prisms are labeled + if they point in the +z direction and – for the opposite.¹⁷ In this scheme, an ordered variant with space group $\text{Imm}2$ and a tripled unit cell was missing.^{16,21} A slight drawback is that the theoretical stacking sequences are not necessarily unambiguous. For the tripled unit cell, six prisms have to be aligned, for which two sequences are possible: + + – – + +, + + – – + + or – + – – + –, – + – – + –. The difference lies only in converting the first prism from + to –, which means an exchange of gold and tin atoms for that prism. In the present paper we report on the structure determination of YbAuSn , which turned out to be the *missing link*. YbAuSn has already been reported in the literature,^{24,25} but the wrong structure type has been assigned. On the basis of X-ray powder data it was concluded that YbAuSn is isotypic with the TiNiSi type. In these investigations, the weak superstructure reflections have most likely

been overlooked. In addition to the precise structure determination, we report also on magnetic susceptibility, electrical resistivity, and ^{119}Sn Mössbauer spectroscopic measurements.

Experimental Procedures

Synthesis. Starting materials for the syntheses of YbAuSn were ingots of ytterbium (Heraeus), gold wire (\varnothing 1 mm, Degussa), and a tin bar (Heraeus), all with stated purities better than 99.9%. Homogeneous samples of YbAuSn were obtained via two different synthesis techniques. In the first experiment the elements (ideal 1:1:1 starting composition) were reacted in glassy carbon crucibles (Sigradur G, glassy carbon, type GAZ006) under flowing argon in a high-frequency furnace (Hüttinger Elektronik, Freiburg, Typ TIG 1.5/300). The argon was purified over silica gel, molecular sieves, and titanium sponge (900 K). The glassy carbon crucibles were placed in a water-cooled sample chamber as described in detail elsewhere.¹⁶ During the inductive heating process, gold and tin first form an alloy at low temperature that finally reacts with the ytterbium pieces in a strongly exothermic reaction. To ensure homogeneity and good crystallization, the sample was held for 2 h at ~ 800 K. No weight losses were observed for these reactions. After cooling to room temperature, the sample could easily be separated from the glassy carbon crucible by pounding at its base.

Alternatively, the elemental components were reacted in a sealed tantalum tube²⁶ in a quartz glass sample chamber of a high-frequency furnace (Kontron Roto-Melt, 1.2 kW) under flowing argon.²⁷ The tube was first heated with the maximum power output (about 1700 K) until the reaction started, which was indirectly visible by a heat flash. Then the sample was held at about 1400 K for another 2–3 min, followed by quenching. To ensure good crystallinity and to further check for possible reconstructive phase transitions, two more samples in tantalum tubes were prepared following the above syntheses conditions, but additionally the tantalum tubes were sealed in evacuated silica tubes and annealed for 10 days at 1170 K in conventional resistance furnaces. The second sample was kept for another 35 days at 870 K. After the annealing procedures the samples were cooled to room temperature by radiative heat loss. The samples could easily be separated from the tantalum tubes by mechanical fragmentation.

In both synthesis techniques no reactions of the samples with the crucible material could be detected. Polycrystalline samples and powders of YbAuSn are stable in moist air. No deterioration was observed after several months. Single crystals exhibit metallic luster and powders are dark gray.

The samples were routinely investigated by EDX analyses using a Leica 420 I scanning electron microscope with YbF_3 , gold, and tin as standards. The compositions were all close to the ideal value. No impurity elements heavier than sodium ($Z = 11$) could be detected.

Structural Characterization. The samples were routinely characterized through Guinier powder patterns or powder diffractograms (Stoe Stadi P) using $\text{Cu K}\alpha 1$ radiation and α -quartz ($a = 491.30$ pm, $c = 540.46$ pm) as an internal standard. The strongest reflections could be indexed hexagonally according to the AlB_2 type with $a = 470.0$ and $c = 365.2$ pm. Peak broadening and additional medium strong reflections led to an orthorhombic indexing with the lattice parameters a , $2c$, and $\sim a\sqrt{3}$, satisfying the extinction conditions of a body-centered lattice. The latter corresponds to a KHg_2 ²⁸ type subcell (space group Imma) with $a' = 469.7$, $b' = 730.4$, and $c' = 812.7$ pm with an $a'\sqrt{3}/c'$ ratio of 1.001, very close to the ideal value of an orthohexagonal setting. Additional very weak reflections, confirmed by the single crystal work to belong to YbAuSn (see

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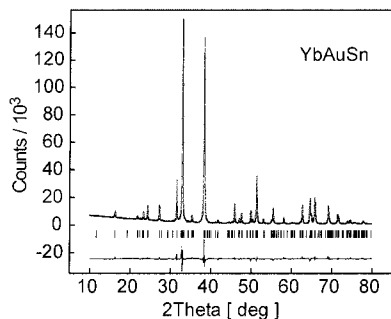
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Table 1. Crystal Data and Structure Refinement for YbAuSn

| | |
|--|---|
| empirical formula | YbAuSn |
| molar mass | 488.7 g/mol |
| lattice parameters (Guinier data) | $a = 469.7(1)$ pm $b = 2191.2(4)$ pm $c = 812.7(1)$ pm $V = 0.8364(2)$ nm ³ |
| formula units per cell | $Z = 12$ |
| space group | <i>Imm2</i> (No. 44) |
| calculated density | 11.64 g/cm ³ |
| crystal size | 20 × 30 × 40 μm ³ |
| transmission (max: min) | 3.91 |
| absorption coefficient | 94.3 mm ⁻¹ |
| F(000) | 2388 |
| θ range | 2–35° |
| range in <i>hkl</i> | ±7, ±35, ±13 |
| total no. reflections | 8068 |
| independent reflections | 2061 ($R_{\text{int}} = 0.1099$) |
| reflections with $I > 2\sigma(I)$ | 1124 ($R_{\sigma} = 0.0718$) |
| data/restraints/parameters | 2061/2/58 |
| goodness-of-fit on F^2 | 0.974 |
| final R indices [$I > 2\sigma(I)$] | $R1 = 0.0449$; $wR2 = 0.0895$ |
| indices (all data) | $R1 = 0.1098$; $wR2 = 0.1114$ |
| R indices (subcell reflections) ^a | 616 > $2\sigma(I)$, $R = 0.0282$ |
| R indices (supercell reflections) ^a | 508 > $2\sigma(I)$, $R = 0.1051$ |
| extinction coefficient | 0.00036(3) |
| absolute structure parameter | 0.00(3) |
| largest diff peak and hole | 3.81 and -4.11 e/Å ³ |

^a These values were calculated with the program RWERT⁴⁷ using the formula $R = \sum |F_o - F_c| / \sum |F_o|$.

**Figure 1.** Powder diffractometer measurement and Rietveld refinement of YbAuSn (Cu K α 1 radiation).

below), forced a tripling of the subcell b axis, resulting in the lattice parameters listed in Table 1. To ensure correct indexing, the observed patterns were compared with a calculated one,²⁹ taking the positions of the refined structure. Powder diffractograms were taken of three samples: the 1400 K high-temperature sample as well as the at 1170 and 870 K annealed samples. Rietveld refinements with the FULLPROF program³⁰ using the positional parameters of the single crystal work for all three diffractograms revealed basically the same results and were in accordance with the single crystal work, except for larger standard deviations: YbAuSn derives from the aristotype AlB₂ and adopts a superstructure of the KHg₂ type with a tripled b axis. This is at least verified for the examined temperature range from 800 up to 1400 K, but it can be assumed that YbAuSn adopts the found superstructure over the whole temperature range. Exemplary, the Rietveld fit of the powder pattern of the 1170 K sample is shown in Figure 1.

Several small, irregularly shaped single crystals of YbAuSn could be isolated from the annealed samples by mechanical fragmentation. They were intensively investigated on a Buerger

er precession camera in order to establish both symmetry and suitability for intensity data collection. A set of layer photographs was taken from three crystals originating from the 800, 1170, and 1400 K sample.

Intensity data were recorded at room temperature by use of a four-circle diffractometer (CAD4) with graphite-monochromatized Mo K α radiation ($\lambda = 71.073$ pm) and a scintillation counter with pulse-height discrimination. The scans were taken in the $\omega/2\theta$ mode and empirical absorption corrections were applied on the basis of ψ -scan data.

It turned out that two crystals (800 and 1400 K) consisted of trilling domains, as could be expected from the $\sqrt{3}$ ratio already determined from the powder patterns. Not being aware of the presence of trillings misleads one to falsely double lattice parameters ($2a$ and $2c$) and acknowledge nonspecific space group extinctions (i.e., a mainly and systematically empty reciprocal space). Some layers, however, still show nonspecific extinctions conditions, but they do not translate. This pronounced pseudosymmetry is characteristic of all KHg₂ superstructures. Doubling of a and c can safely be ruled out by looking at the subgroups of *Imma*. No such symmetry reduction is allowed for a space group with an I lattice. The only possible space group satisfying the observed extinction conditions is *Imm2*, which is nicely in accordance with the predicted one.

Initially the structure was solved for a trilling with space group *Imm2* and the atomic positions deduced for the missing link. The correct stacking sequence for YbAuSn was found to be the aforementioned second choice: $- + - - + -$, $- + - - + -$. The refinement unfortunately was clearly hampered by large correlations of the anisotropic displacement parameters and the extra hindrance that, along with the trilling domains, twinning by inversion obviously had to be accounted for. The occurrence of twins of trillings (i.e. six domains with the refined ratios of 1:0.254:0.248:0.118:0.105:0.087) is understandable by the group-subgroup^{31,32} relations (ref 21 and Figure 2), which contains two *translationengleiche* symmetry reductions of index three (t3) and of index two (t2). This refinement resulted in $R1 = 0.094$, $wR2 = 0.179$, 9205 F^2 values ($0\sigma_{F_o}$), and 63 parameters; $R1 = 0.049$ for 5694 F^2 values ($4\sigma_{F_o}$).

At the end an obviously single-domain crystal of YbAuSn was found in the 1170 K sample. Data collections and refinements were carried out on a step by step basis given by the group-subgroup relations presented in Figure 2. To check for possible trilling reflections the strongest superstructure reflections of YbAuSn were calculated taking the parameters of the above refinement. No measurable intensities could be observed at those positions in reciprocal space corresponding to trilling reflections. Therefore, at first a complete data set was collected only for the KHg₂ type subcell (*Imma*, $a = 470$, $b = 730$, and $c = 812.7$ pm). This refinement resulted in $R1 = 0.029$ (352 F^2 and 12 parameters) with the expected mixed Au/Sn occupancy. The refined parameters are used in Figure 2. Following the group-subgroup relations the next refinement was carried out based on the superstructure of the EuAuGe type²² (space group *Imm2*), which would principally allow ordering of gold and tin atoms. As is shown in Figure 2, ordering of gold and tin was not possible. Instead mixed occupancies of 0.66 Au/0.33 Sn and 0.33 Au/0.66 Sn were found for the two positions ($R1 = 0.024$, 386 F^2 , and 23 parameters). The latter result could also be expected, because the weak superstructure reflections have purposely been ignored.

In the final step a complete data set including the very weak superstructure reflections causing the b axis to be tripled was collected by employing a longer counting time (180 s per reflection) to ensure a reasonable good counting statistic.

The starting atomic parameters were assigned on the basis of the precalculated ones, and the structure was successfully

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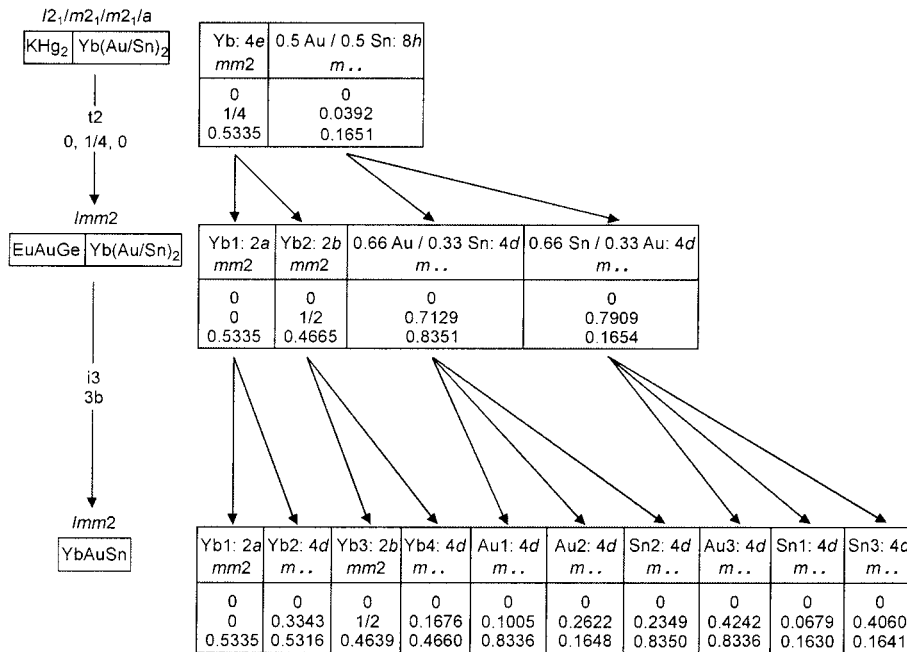


Figure 2. Bärnighausen tree for the group-subgroup relations of the KHg_2 type subcell of $\text{Yb}(\text{Au}/\text{Sn})_2$ and ordered YbAuSn via the EuAuGe type. The evolution of the atomic sites together with the index of the *translationengleiche* (t) and the isomorphic (i) symmetry reduction is given as well as the origin shift.

Table 2. Atomic Coordinates and Isotropic Displacement Parameters for YbAuSn

| atom | Wyckoff position | x | y | z | U_{eq}^a |
|------|------------------|---|------------|-----------|-------------------|
| Yb1 | 2a | 0 | 0 | 0.5335(5) | 140(7) |
| Yb2 | 4d | 0 | 0.33430(9) | 0.5316(3) | 108(5) |
| Yb3 | 2b | 0 | 1/2 | 0.4639(5) | 123(6) |
| Yb4 | 4d | 0 | 0.16763(9) | 0.4660(3) | 123(5) |
| Au1 | 4d | 0 | 0.10052(9) | 0.8336(4) | 113(4) |
| Au2 | 4d | 0 | 0.2622(1) | 0.1648(3) | 143(4) |
| Au3 | 4d | 0 | 0.42420(9) | 0.8336(4) | 117(4) |
| Sn1 | 4d | 0 | 0.0679(2) | 0.1630(6) | 98(6) |
| Sn2 | 4d | 0 | 0.2349(2) | 0.8350(5) | 82(4) |
| Sn3 | 4d | 0 | 0.4060(2) | 0.1641(6) | 179(8) |

^a U_{eq} (pm^2) is defined as one-third of the trace of the orthogonalized U_{ij} tensor.

refined with anisotropic displacement parameters using SHELXL-97³³ (full-matrix least-squares on F^2). The refinements readily converged to the residuals listed in Table 1 and a subsequent difference Fourier synthesis revealed no significant residual peaks. Especially the Flack^{34,35} parameter gave no indication of twinning by inversion, which is possible due to the translationengleiche symmetry reduction of index two (t2) in going from the KHg_2 to the EuAuGe type. All relevant crystallographic details are listed in Table 1. Atomic coordinates and interatomic distances are given in Tables 2 and 3. Listings of the anisotropic displacement parameters and the structure factors are available.³⁶

As a check for the correct compositions, the occupancy parameters were varied in a separate series of least-squares cycles along with the displacement parameters. All sites were fully occupied within two to five standard deviations, and in the final cycles the ideal composition was assumed. The X-ray data gave no hint for gold/tin mixing within the polyanion.

Physical Property Investigations. The magnetic susceptibilities of polycrystalline pieces of YbAuSn were deter-

Table 3. Interatomic Distances (pm) in the Structure of YbAuSn^a

| | | | | | | | |
|------|---|-----|-------|------|---|-----|-------|
| Yb1: | 2 | Au1 | 328.6 | Yb3: | 2 | Sn3 | 319.1 |
| | 4 | Sn3 | 329.9 | | 4 | Sn1 | 321.7 |
| | 4 | Au3 | 330.3 | | 4 | Au1 | 338.9 |
| | 2 | Sn1 | 335.9 | | 2 | Au3 | 343.4 |
| | 2 | Yb4 | 371.4 | | 2 | Yb2 | 367.2 |
| | 2 | Yb3 | 421.3 | | 2 | Yb1 | 421.3 |
| Yb2: | 1 | Au3 | 314.8 | Yb4: | 1 | Au2 | 320.8 |
| | 2 | Au1 | 318.5 | | 2 | Au2 | 323.9 |
| | 2 | Sn2 | 322.0 | | 2 | Sn3 | 327.2 |
| | 1 | Sn2 | 329.1 | | 2 | Au3 | 327.4 |
| | 2 | Au2 | 334.1 | | 1 | Sn1 | 329.2 |
| | 2 | Sn1 | 335.4 | | 1 | Au1 | 333.0 |
| | 1 | Au2 | 337.3 | | 1 | Sn2 | 334.2 |
| | 1 | Sn3 | 337.4 | | 2 | Sn2 | 334.8 |
| | 1 | Yb3 | 367.2 | | 1 | Yb2 | 369.1 |
| | 1 | Yb4 | 369.1 | | 1 | Yb1 | 371.4 |
| | 2 | Yb4 | 424.1 | | 2 | Yb2 | 424.1 |
| Au1: | 2 | Sn3 | 272.6 | Sn1: | 2 | Au3 | 273.3 |
| | 1 | Sn1 | 277.1 | | 1 | Au1 | 277.1 |
| | 1 | Sn2 | 294.4 | | 1 | Sn1 | 297.6 |
| | 2 | Yb2 | 318.5 | | 2 | Yb3 | 321.7 |
| | 1 | Yb1 | 328.6 | | 1 | Yb4 | 329.2 |
| | 1 | Yb4 | 333.0 | | 2 | Yb2 | 335.4 |
| | 2 | Yb3 | 338.9 | | 1 | Yb1 | 335.9 |
| Au2: | 2 | Sn2 | 272.6 | Sn2: | 2 | Au2 | 272.6 |
| | 1 | Sn2 | 274.6 | | 1 | Au2 | 274.6 |
| | 1 | Sn3 | 315.0 | | 1 | Au1 | 294.4 |
| | 1 | Yb4 | 320.8 | | 2 | Yb2 | 322.0 |
| | 2 | Yb4 | 323.9 | | 1 | Yb2 | 329.1 |
| | 2 | Yb2 | 334.1 | | 1 | Yb4 | 334.2 |
| | 1 | Yb2 | 337.3 | | 2 | Yb4 | 334.8 |
| Au3: | 1 | Sn3 | 271.5 | Sn3: | 1 | Au3 | 271.5 |
| | 2 | Sn1 | 273.3 | | 2 | Au1 | 272.6 |
| | 1 | Yb2 | 314.8 | | 1 | Au2 | 315.0 |
| | 2 | Yb4 | 327.4 | | 1 | Yb3 | 319.1 |
| | 2 | Yb1 | 330.3 | | 2 | Yb4 | 327.2 |
| | 1 | Au3 | 332.2 | | 2 | Yb1 | 329.9 |
| | 1 | Yb3 | 343.4 | | 1 | Yb2 | 337.4 |

^a All distances of the first coordination spheres are listed. Standard deviations are all equal or less than 0.4 pm.

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mined with a SQUID magnetometer (Quantum Design, Inc.) in the temperature range from 4.2 to 300 K with magnetic flux densities up to 5.5 T. The specific resistivities were

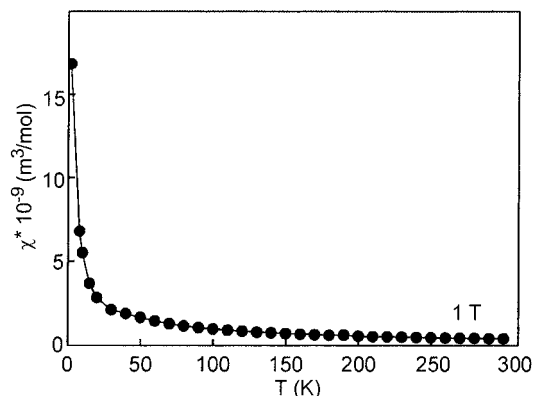


Figure 3. Temperature dependence of the magnetic susceptibility of YbAuSn measured at a flux density of 1 T.

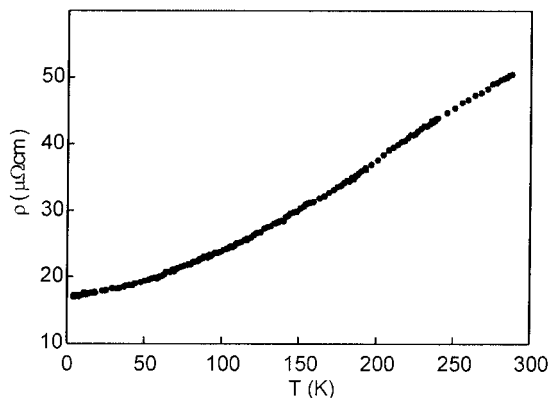


Figure 4. Temperature dependence of the specific resistivity of YbAuSn.

measured on small, irregularly shaped polycrystalline blocks (typical dimensions $1 \times 1 \times 2 \text{ mm}^3$) using a four-probe technique. Cooling and heating curves measured between 4.2 and 300 K were identical within the error limits, also for different samples. ^{119}Sn Mössbauer spectroscopic experiments were performed with a $\text{Ca}^{119\text{m}}\text{SnO}_3$ source at 300 K on the same polycrystalline samples used for the magnetic and electrical measurements. A palladium foil of 0.05 mm thickness was used to reduce the thin K X-rays concurrently emitted by the source.

Results and Discussion

Magnetic Properties. The temperature dependence of the magnetic susceptibility of YbAuSn is displayed in Figure 3. The susceptibilities were only weakly dependent on the external field, indicating only very small amounts of ferromagnetic impurities. The 1 T (Figure 3) and 3 T data were practically identical. Down to about 50 K the susceptibilities are very small and nearly independent of temperature. The upturn below 50 K may be attributed to a minor degree of paramagnetic impurities, although our Guinier diagrams showed single-phase YbAuSn. The susceptibility at 300 K is $0.4(1) \times 10^{-9} \text{ m}^3/\text{mol}$, indicating Pauli paramagnetism. This is in agreement with the metallic behavior discussed below. Our magnetic data entirely match the previous investigations by Katoh et al.²⁴ and Kaczowski et al.²⁵

Electrical Properties. The temperature dependence of the specific resistivity of YbAuSn is presented in Figure 4. The specific resistivity decreases with decreasing temperature as expected for a metal. The room-

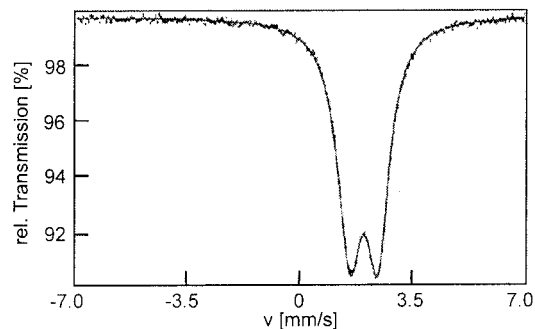


Figure 5. Experimental and simulated ^{119}Sn Mössbauer spectrum of YbAuSn at 4.2 K.

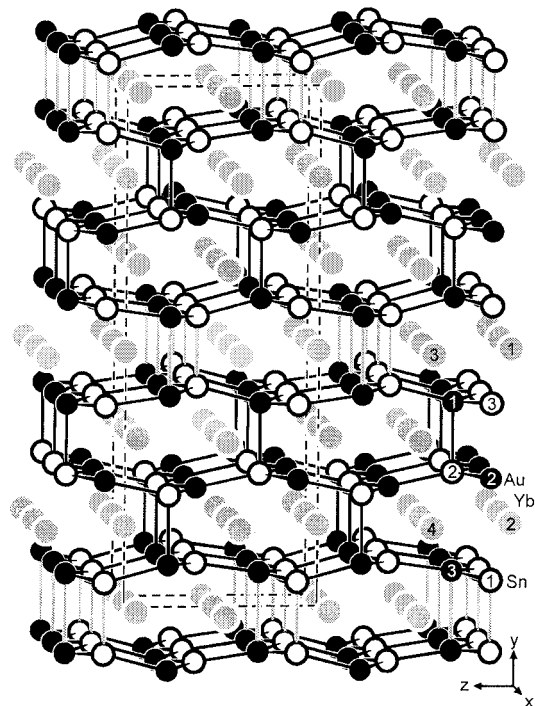


Figure 6. Crystal structure of YbAuSn. Ytterbium, gold, and tin atoms are drawn as gray, black filled, and open circles, respectively. Single digits correspond to the atom designations. Along with the unit cell the three-dimensional [AuSn] polyanion is outlined. The different bonding interactions within the hexagons and between the stacked layers are emphasized by black and gray lines. For details, see the text.

temperature value is $50 \pm 10 \mu\Omega \text{ cm}$. The relatively large error limit accounts for the different values obtained for several irregularly shaped samples. At 4.2 K the specific resistivity has dropped to $17 \pm 5 \mu\Omega \text{ cm}$, resulting in a resistivity ratio of $\rho(4.2 \text{ K})/\rho(300 \text{ K}) = 0.34$. Over the whole temperature range, no anomaly could be observed. The shape of the resistivity curve presented by Katoh et al.²⁴ is slightly different. This is most likely due to the polycrystalline character of both samples.

^{119}Sn Mössbauer Spectroscopy. The ^{119}Sn Mössbauer spectrum at 4.2 K of YbAuSn is shown in Figure 5 together with a theoretical transmission integral fit. Despite the three crystallographically different tin sites, the spectrum could be fit with one signal at an isomer shift of $\delta = 1.922(5) \text{ mm/s}$ subjected to quadrupole splitting of $\Delta E_Q = 0.929(2) \text{ mm/s}$. The experimental line width is $\Gamma = 1.05(5) \text{ mm/s}$. Since the coordinations of the Sn1, Sn2, and Sn3 atoms are quite similar, the

longer average gold–tin distances found in NaAuSn (275)⁴³ and EuAuSn (278). One can safely assume strong covalent bonding within the hexagons that extend two-dimensionally in YbAuSn. The weaker interlayer gold–tin bonding interactions connecting the layers are quite longer (294 and 315 pm). The tin–tin interactions, however, are even shorter (298 pm) than those found in β -tin,⁴⁴ i.e., 4×302 pm and 2×318 pm in the form of a flattened tetrahedron with two additional tin atoms. The gold–gold contacts of 332 pm are in the range found for EuAuSn and CaAuSn, 335 and 326 pm, respectively, whereas they are shorter in EuAuGe (316 pm). In the latter compounds one surely has to consider the large relativistic contraction of the gold 6s orbitals¹⁸ leading to gold–gold interactions.

Depending on the elemental combinations of the equiatomic AAuX compounds, A = electropositive element and X = main group element, one encounters different electronic situations paired with space requirements. The nature of the A and X elements influences the s electron density at the gold atoms and triggers consequently the various superstructures. The global minimum on the energy hypersurface is determined by

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an arrangement given by the KHg₂ type. Introducing secondary gold–gold and in the case of YbAuSn tin–tin bonding fine-tunes the structure. The latter can be held responsible for the formation of the superstructure of YbAuSn. The reason for the formation of the other KHg₂ type superstructures such as CaCuGe,¹⁶ YPdSi,⁴⁵ and UPtGe⁴⁶ may be ascribed to similar secondary bonding interactions.

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